



LOW TEMPERATURE DETECTORS

Radioactive background reduction in the Milano neutrino mass experiment





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A neutrino mass experiment studies the end-point of a beta decay spectrum: since the event rate in this region is extremely low the radioactive background must be as low as possible. Moreover in a calorimetric neutrino mass experiment the β source cannot be switched off, therefore the unknown radioactive background below the β spectrum must be minimized to avoid deformations of the spectrum. In this poster we discuss the radioactive background observed in the Milano neutrino mass experiment.

1. Experimental Configurations

Material selection in detector proximities • Roman Pb with lowest ²¹⁰Pb content ($A_{210Pb} \leq 4 \text{ mBq/kg}$) • Pb-free solder (Sn/Ag alloy)

4. RUN 9 results

- high background continuum due to ⁵⁵Fe Inner-Bremsstrahlung (IB) ► IB gammas excite Pb L and K X-rays in Pb shielding close to detectors \blacktriangleright X-rays undergo photoelectric absorption in AgReO₄ mainly on K and L Re shells

RUN 9 (2354 h×mg)

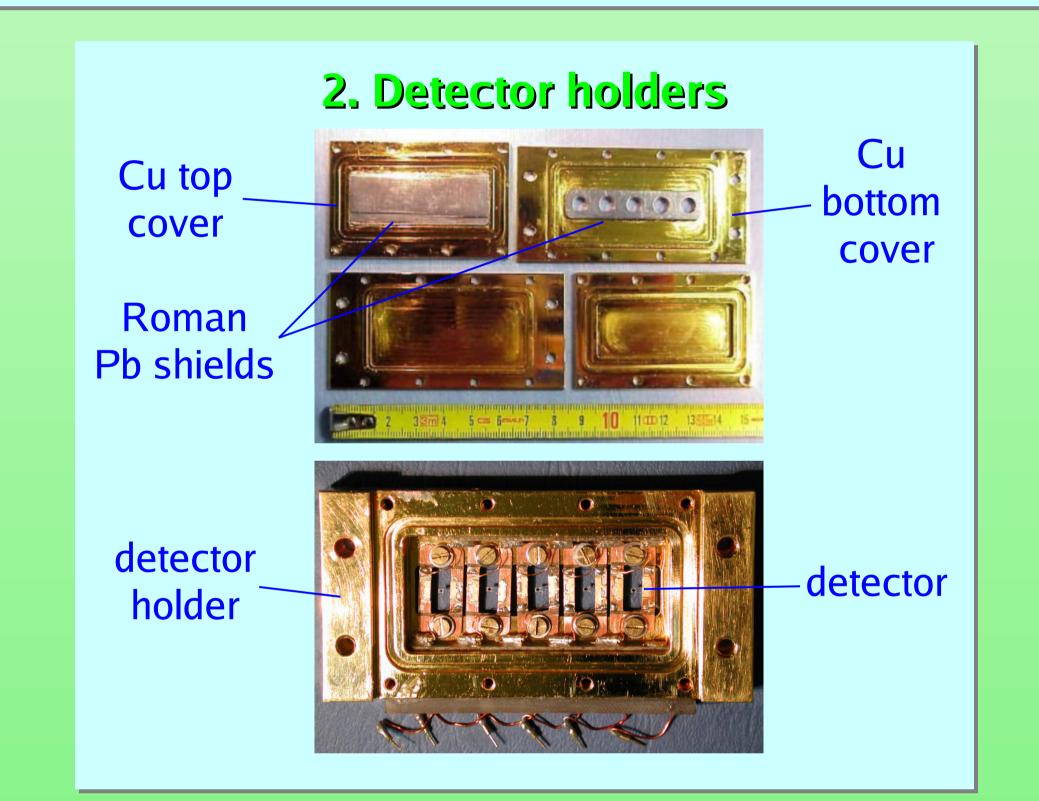
- detectors shielded with 3 mm Roman Pb from above
- two 10 mCi ⁵⁵Fe sources in fixed positions below the detectors shielded by $\leq 1 \text{ cm Cu (top)} + \sim 1 \text{ mm Pb/0.2 mm Cu (shutter)}$
- ▶ 12 kBq ⁵⁵Fe *Inner-Bremsstrahlung* (Q_{IR} = 232 keV)

RUN 10 (1538 $h \times mg$)

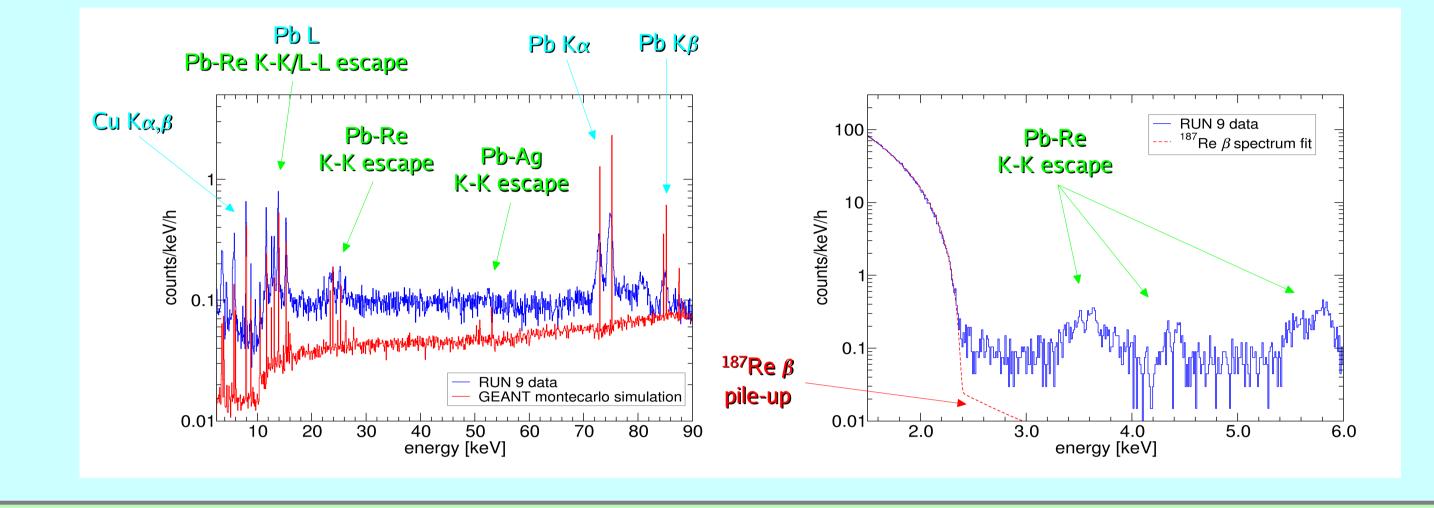
- 5 detectors shielded with Roman Pb (3 mm above, 2 mm below)
- 5 detectors without Pb shield
- without ⁵⁵Fe sources

RUN 14 (8751 h×mg)

- detectors shielded only with Cu
- two 10 mCi ⁵⁵Fe sources which retracts into a ~2 cm thick Roman Pb shield

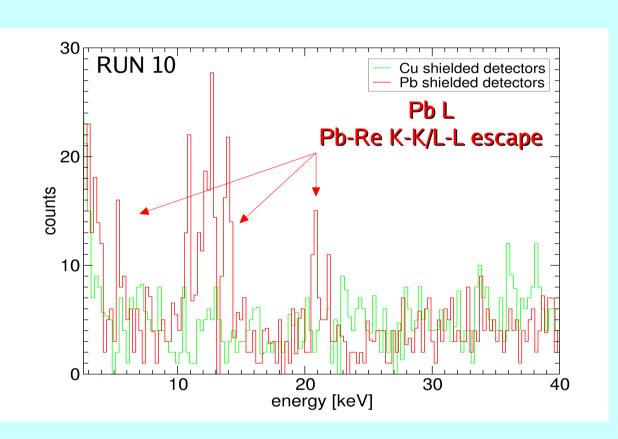


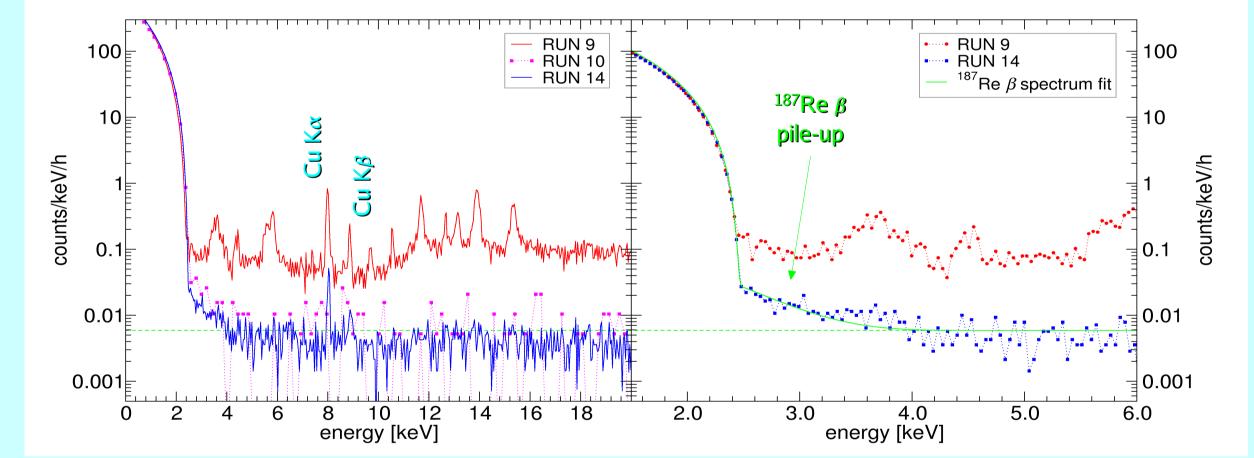
escaping Re K and L X-rays cause escape peaks at low energies • $b \approx 6 \times 10^{-2}$ counts/keV/h



5. RUN 10 and 14 results

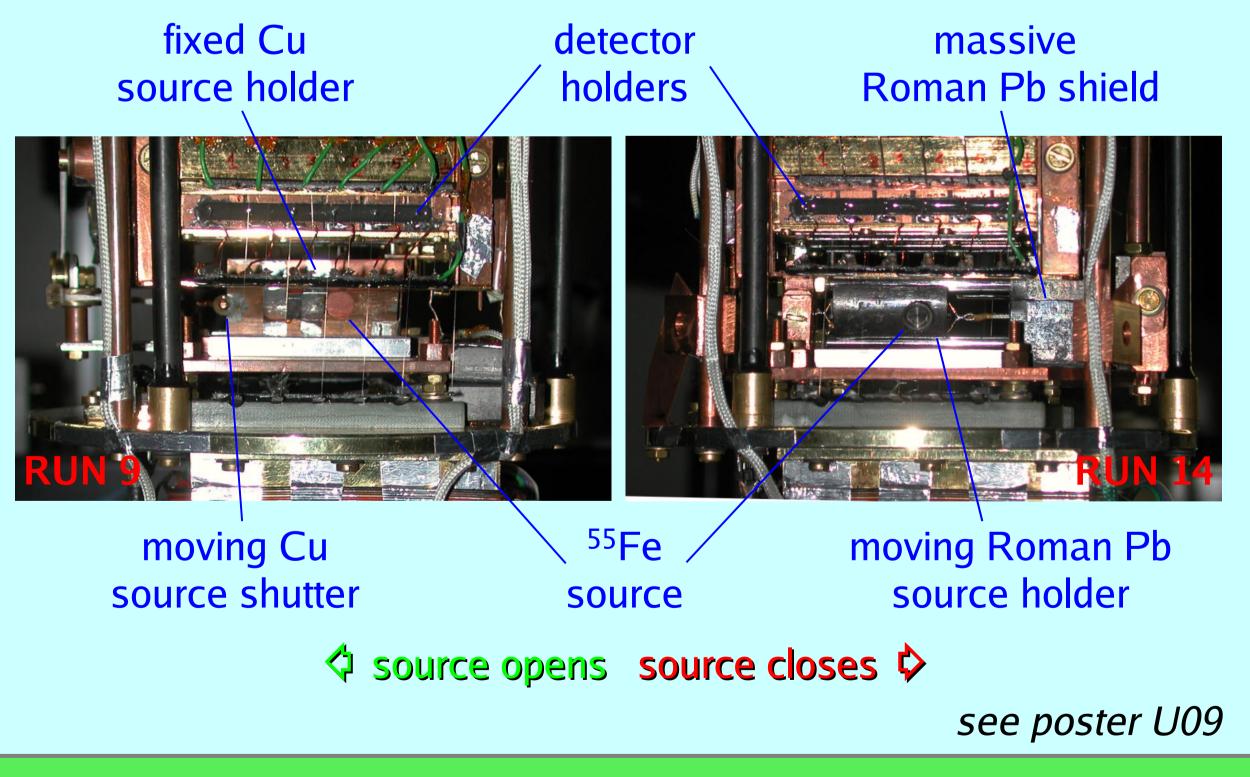
- detector Pb shielding increases background
- shielding the ⁵⁵Fe IB removes most of the background
- RUN 14 source configuration is equivalent to RUN 10 without ⁵⁵Fe sources
- RUN 14 continuum is constant up to 20 keV and is mainly due by cosmics and environmental radioactivity **b** \approx 7×10⁻³ counts/keV/h
- ► signal-to-noise $S/N(Q_B:Q_B-30 \text{ eV}) = \int N_B(E) / \sqrt{\int b} \approx 12$





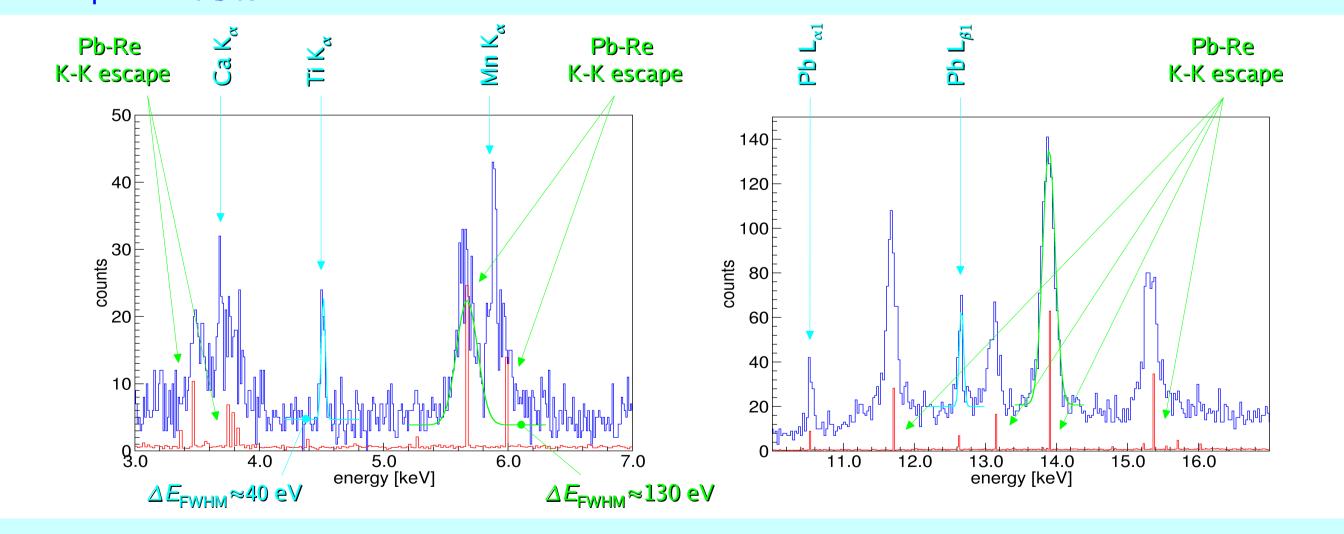


3. ⁵⁵Fe source in RUN 9 and RUN 14



6. Details of the X-ray escape peaks in RUN 9 spectra

escape peaks are broader than X-ray peaks due to intrinsic Pb and Re X-ray line widths ► $Γ_{\text{K-K escape}} = Γ_{\text{Pb K}} + Γ_{\text{Re K}} \approx 60.2 \text{ eV} + 42.1 \text{ eV} = 102 \text{ eV}$



7. Conclusions

- low and constant background is achieved by carefully shielding the ⁵⁵Fe sources
- detectors must be shielded only with low Z materials
- S/N should improve increasing detector size m ($S \propto A_{\beta} \propto m$ and $N \propto \sqrt{b} \propto \sqrt{m}$)
- background can be further reduced by material selection
- cosmics reduction can be attempted by using an active veto shield just above the detectors (p.e. a Si photodiode)

